

IMPACT OF DIFFERENT TYPES OF GRAVITY WAVES ON THE TURBULENT EDDIES, EXCHANGE COEFFICIENTS AND THE SPECTRAL GAP.



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1. INTRODUCTION

Turbulent motions developed in stably stratified atmospheric boundary layers with gravity wave activity show distinctive features compared with turbulence developed in unstable or weakly stable conditions. The flow can show a strong non-stationary behaviour due to the strong stratification, the presence of waves and their eventual collapse into turbulence. The wave-modified turbulence can extract or transfer energy to the wave; the turbulent fluxes usually show a periodic component, and appear sometimes accompanied by counter-gradient transfers produced by motions at the scale of the wave. In this study we compare the effect of three different gravity wave events on the main characteristics of turbulent motions.

2. <u>DATA</u>

- SABLES2006 field campaign was carried out on a relatively flat and homogeneous terrain over an extensive high plain in the northern part of the Iberian Peninsula^[1] (See poster A335, this session, for site characteristics).
- Main instrumentation, installed on a 100-m mast, included wind vanes, cup and sonic anemometers, thermometers and three microbarometers at different levels (plus 3 additional microbarometers deployed at surface level on a triangular array around the site). In this study turbulence is studied from sonic anemometer data at z=19.6 m and z=96.6 m.

3. <u>METHOD: MULTIRESOLUTION (MR) FLUX DECOMPOSITION</u>

fluxes and studying the contribution by different timescales $\tau^{[2]}$.

• In stable conditions, MR cospectra of vertical heat flux ($\langle w \theta \rangle_{\tau}$) is usually studied.

(see adjacent figure).

4. TYPES OF GRAVITY WAVES AT SABLES 2006



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Implications in the evaluation of exchange coefficients: -

- A proper evaluation of turbulent fluxes $\langle u'w' \rangle$ and $\langle w'\theta' \rangle$ (integrating MR cospectra up to the spectral gap) is crucial.
- As a general result, waves overlap, to a greater or lesser extent, with the spectra of turbulent motions.

• If no spectral gap can be identified, uncertainty in the evaluation of $\langle u'w' \rangle$ and $\langle w'\theta' \rangle$ increases (and thus, in K_m and K_h). This effect is more important in strong stability conditions where waves develop, because turbulent fluxes tend to be very small).

6. REFERENCES

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→ K _m = - < u'	$\mathbf{w'} > / \frac{\partial \overline{\mathbf{u}}}{\partial z}$	$\mathbf{K}_{\mathbf{h}} =$	$- < \theta' w' > / \frac{\partial \theta}{\partial z}$

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